

## Optical turbulence profiling at White Sands Missile Range North Oscura Peak

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November 2, 2006

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## Introduction

• AFRL Starfire Optical Range owns/operates a device known as the "Differential-Tilt Turbulence Profiler"

#### Development timeline

- O 2000: AFRL/DE and AFOSR developed original concept, theory, initial design
- O 2001-2003: AFRL/DE and MZA integrated the laser & optical system, developed data acquisition software
- O 2002-2005: AFRL/DE conducted simulations, developed processing algorithms, initial testing and data analysis
- O 2006: Full-scale system test at North Oscura Peak (NOP)

#### • <u>Testing timeline</u>

- 2003: NOP to Bug (~1 km), NOP to Beck (~10 km)
  - Initial data, but issue with source assembly
- O 2004: SOR to 2-mile (~3 km)
  - Revised source assembly
  - Reasonable test data and profiling results for 3 km path
- O 2005: NOP to Salinas (~50 km)
  - Poor SNR limited measurements, profiling questionable
- O 2006: NOP to Beck (~10 km)
  - Good SNR, reasonable profiles, large data volume, other diagnostics



## **Differential-Tilt Turbulence Profiler**



MRW - 11/2/2006



## Estimation of Coherence Diameter from Profiler Centroid Data



- Ap-1 / Ap-2 with  $s/D = 1.5875 \rightarrow$
- Ap-3 / Ap-4 with  $s/D = 6.9850 \rightarrow$

• Ap-5 / Ap-6 with 
$$s/D = 1.5875 \rightarrow$$

- Can use centroid (tilt) data directly from profiler units with standard techniques for r0 estimation
  - **O** Estimation from tilt-variance:

$$\sigma_T^2 = 0.1816 D^{-1/3} \lambda^2 r 0^{-5/3}$$
$$r_0 = \left(\frac{\sigma_T^2 D^{1/3}}{0.1816 \lambda^2}\right)^{-3/5}$$

• Differential-tilt variance for 3 aperture pairs on each unit

$$\sigma_{\Delta T}^2 = 0.1943 D^{-1/3} \lambda^2 r 0^{-5/3}$$
$$\sigma_{\Delta T}^2 = 0.2619 D^{-1/3} \lambda^2 r 0^{-5/3}$$
$$\sigma_{\Delta T}^2 = 0.1943 D^{-1/3} \lambda^2 r 0^{-5/3}$$



### • Profiler theory published in SPIE proceedings

- Whiteley, M. R., Washburn, D. C., and Wright, L. A., "Differential-tilt technique for saturation-resistant profiling of atmospheric turbulence," *SPIE Proceedings on Adaptive Optical Systems Technology II* 4494, (2001).
- Difference of differential-tilt variances define measurement set that can be related to turbulence distribution over path

continuous: 
$$m_k = \int_0^1 d\xi \ C_n^2(\xi L) \ w_{\delta k}(\xi)$$
  
discrete:  $m_k = \sum_{i=1}^N C_{ni}^2 p_{ki} \to \mathbf{m} = P\mathbf{c}$ 

- Relation can be inverted through appropriate numerical technique
- Constraints can be applied to inversion using r0 estimates for profiler used as differential-image-motion monitor

$$\mathbf{c} = (P^T P)^{-1} (P^T \mathbf{m} + \lambda_A \mathbf{w}_{r_A} + \lambda_B \mathbf{w}_{r_B})$$



## Wave-Optics Simulation of Turbulence Profiling





# **NOP-to-Beck Site Data Collection**

- Turbulence profiler data was collected on NOP-to-Beck path
  - O 9.6 km, ~900 m altitude change
  - O April-August 2006
  - O Other data items
    - Scintillometer
    - Tactical imagery

North Oscura Peak O. Beck Site

- 33 test days with profiler
- 21 days of good quality/volume for profiling
- ~4500 data files
- 726 profiles
  - O 0000-0600: 120
  - O 0600-1200: 86
  - O 1200-1800: 110
  - O 1800-2400: 410





# **Data Quality Checking**

- Data filtering implemented to reduce noise sources in differential-tilt estimation
- Spot SNR
  - O Average pixel SNR > 2 for all subapertures
- Saturated spot image
  - **O** No saturated pixels used in centroid
- Spot clipping
  - Subaperture spot too close to centroid processing boundary
- Telescope jitter
  - O Quantified by averaging shift over all subapertures
  - **O** Retained data with jitter std < 7 pixels
- Required following attributes for profile processing
  - At least 10% of frames in a file must pass all quality checks
  - **O** At least 200 frames total passing quality checks
- SNR was most common reason for data filtering
  - **O** Especially when scintillation was high





# **Profiling Example: Day 172**







# **Comparison with Scintillometer**



- Rytov number from turbulence profiles compared with scintillometer estimate of Rytov number (when available)
- It can be shown that for any turbulence profile, the following inequality applies:  $\mathcal{R} \leq \frac{1}{9.1314} \left(\frac{\lambda}{r_0 \theta_0}\right)^{5/6}$
- A good "rule of thumb" approximation is:
- When scintillometer is in bounds implied by profiler, scintillometer and profiler give consistent Rytov number estimates

 $\mathcal{R}\simeq 15.53^{-1}\left(rac{\lambda}{r_{
m O} heta_{
m O}}
ight)^{5/6}$ 



## Comparison with Standard Turbulence Models

#### "Night" 20:00-06:00 "Morning" 06:00-12:00 "Afternoon" 12:00-20:00 local time 20 hr - 6hr local time 6 hr - 12hr local time 12 hr - 20hr 10<sup>-13</sup> 10<sup>-13</sup> · 10<sup>-13</sup> $10^{-1}$ 10 10 10<sup>-15</sup> 10<sup>-15</sup> 10<sup>-15</sup> $C_n^2 \, (m^{-2/3})$ $C_n^2 (m^{-2/3})$ $C_n^2 \, (m^{-2/3})$ 10 10 10 10<sup>-13</sup> 10<sup>-1</sup> 10 10% (best) 10% (best) 10% (best) 10<sup>-18</sup> 50% (median) 50% (median) 10 10 50% (median) 90% (worst) 90% (worst) 90% (worst) HV57(h<sub>and</sub>) HV57(hand) HV57(h<sub>and</sub>) $10^{-1}$ 10 10 Clear1/Night(h and +1230 Clear1/Night(h<sub>and</sub>+1230) Clear1/Night(hgnd+1230) 10<sup>-20</sup> 10<sup>-20</sup> 10 0.2 0.6 0.2 0.4 0.6 0.8 0.2 0.4 0.6 0.8 0 0.4 0.8 0 1 0 NOP <<< normalized path position >>> Beck NOP <<< normalized path position >>> Beck NOP <<< normalized path position >>> Beck local time 20 hr - 6hr local time 6 hr - 12hr local time 12 hr - 20hi 10 10<sup>1</sup> 10 50% (median 50% (median) 50% (median) $\alpha \times \text{CLEAR-1/Night(}h_{\text{gnd}})$ $\alpha \times CLEAR-1/Night(h_{gnd})$ CLEAR-1/Night(h<sub>gnd</sub>) 10<sup>0</sup> 10<sup>0</sup> 10<sup>0</sup> 10 10 10 х 8 10-2 $10^{-2}$ 10 0.2 0.4 0.6 0.8 0 1 0.2 0.6 0 0.2 0.4 0.6 0.8 0 0.4 0.8 1 NOP <<< normalized path position >>> Beck NOP <<< normalized path position >>> Beck NOP <<< normalized path position >>> Beck



- Differential-tilt turbulence profiler successfully taken from concept to development and into application
  - ~10 km path from North Oscura Peak to valley floor at WSMR
- Substantial test data has been collected and analyzed
  - **O** Filtering for data quality important to assuring turbulence estimate
  - **O** Profile estimates available around the clock
  - O Profiles and derived atmospheric propagation parameters consistent with expected trends
- Profile estimates validated using independent measurements
  - **O** Consistency between profile Rytov and scintillometer
- Diurnal trends (median conditions)
  - Compare favorably to CLEAR-1 model given propagation height above ground
  - **O** Night
    - ◆ < 0.5 x CLEAR-1/Night
  - **O** Morning
    - 0.5 1.0 x CLEAR-1/Night
  - **O** Afternoon
    - End-points (near ground) 2.0 4.0 x CLEAR-1/Night
    - Otherwise, 0.5 1.0 x CLEAR-1/Night